

---

---

# CS680: Radiometry

---

---

**Sung-Eui Yoon**  
(윤성의)

**Course URL:**  
<http://jupiter.kaist.ac.kr/~sungeui/SGA/>

**KAIST**



# Announcements

---

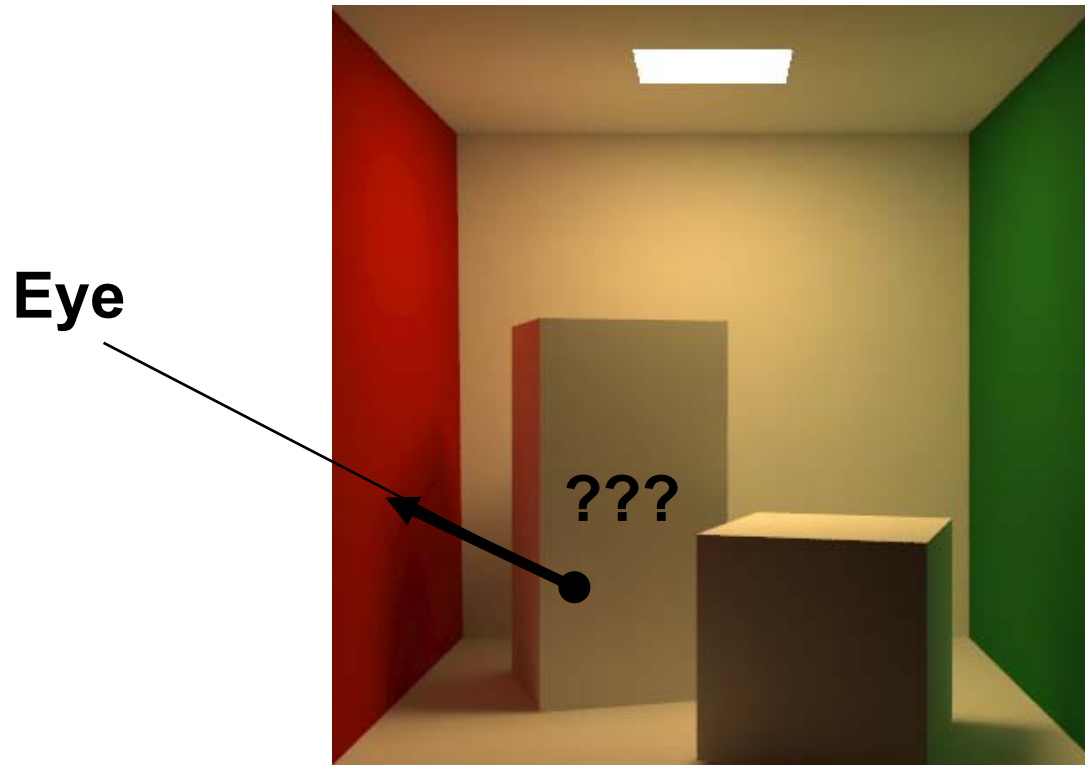
---

- **2 papers for each student**
  - **Choose 4 papers from the paper list**
  - **Send them (titles of 4 papers) to TA (Bochang Moon) by Oct-11 (Mon)**
  - **Look at videos and talk files (captured talk video or presentation files)**
  
- **Schedule of student presentations**
  - **Will be decided on Oct-12 (Tue)**
  - **Presentations will start after the mid-term**

# Motivation

---

---

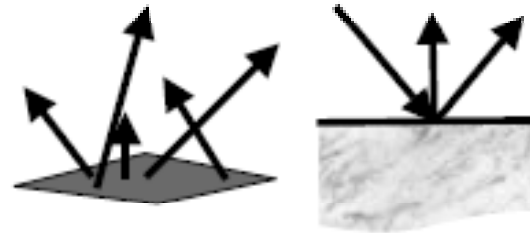


# Light and Material Interactions

---

---

- **Physics of light**
- **Radiometry**
- **Material properties**
  
- **Rendering equation**



From kavita's slides

# Models of Light

---

---

- **Quantum optics**
  - **Fundamental model of the light**
  - **Explain the dual wave-particle nature of light**
- **Wave model**
  - **Simplified quantum optics**
  - **Explains diffraction, interference, and polarization**
- **Geometric optics**
  - **Most commonly used model in CG**
  - **Size of objects  $\gg$  wavelength of light**
  - **Light is emitted, reflected, and transmitted**

# Radiometry

---

---

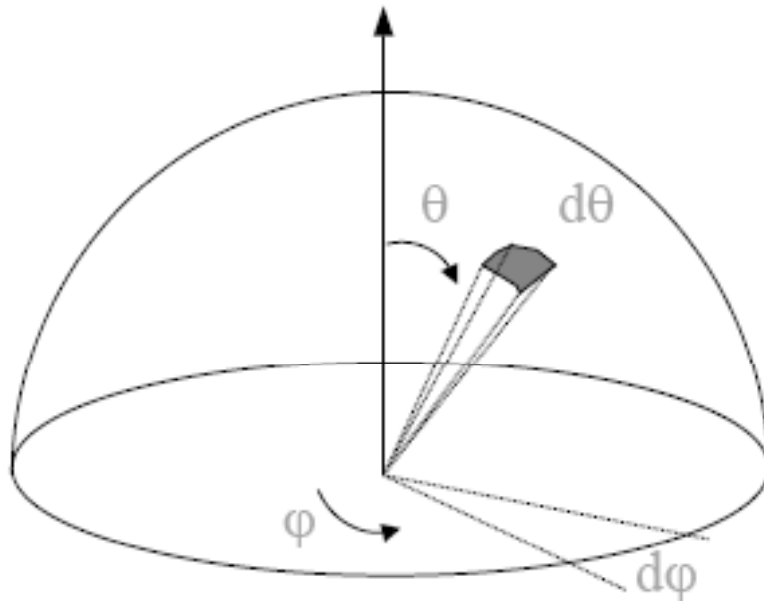
- **Measurement of light energy**
  - **Critical component for photo-realistic rendering**
- **Light energy flows through space**
  - **Varies with time, position, and direction**
- **Radiometric quantities**
  - **Densities of energy at particular places in time, space, and direction**
- **Photometry**
  - **Quantify the perception of light energy**

# Hemispheres

---

---

- Hemisphere
  - Two-dimensional surfaces
- Direction
  - Point on (unit) sphere



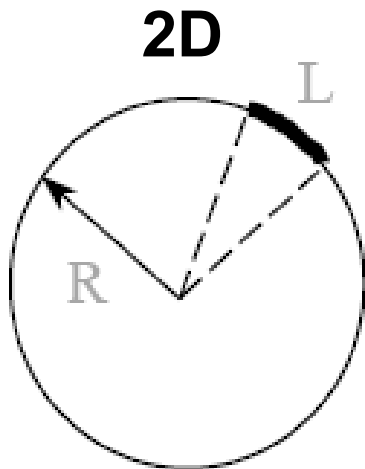
$$\theta \in [0, \frac{\pi}{2}]$$
$$\phi \in [0, 2\pi]$$

From kavita's slides

# Solid Angles

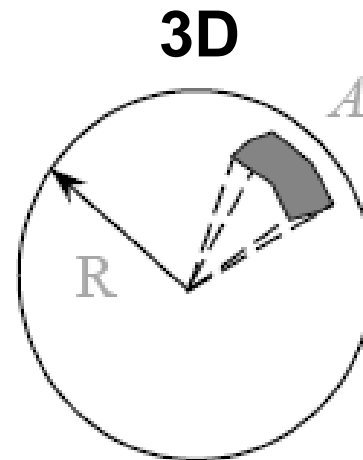
---

---



$$\theta = \frac{L}{R}$$

**Full circle  
=  $2\pi$  radians**



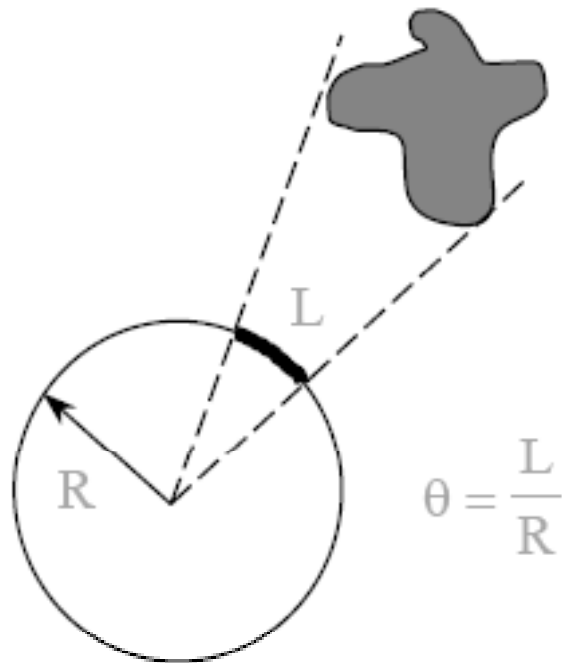
$$\Omega = \frac{A}{R^2}$$

**Full sphere  
=  $4\pi$  steradians**



# Solid Angles

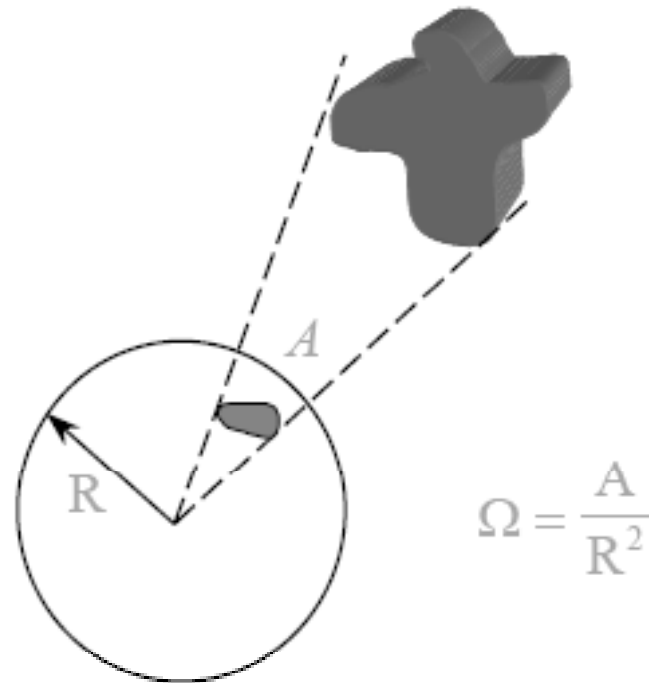
2D



$$\theta = \frac{L}{R}$$

**Full circle  
=  $2\pi$  radians**

3D

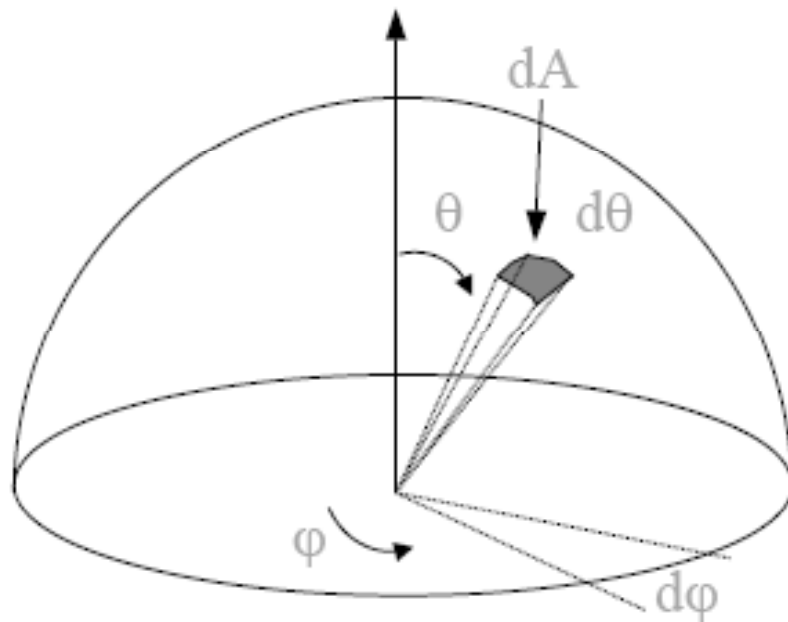


$$\Omega = \frac{A}{R^2}$$

**Full sphere  
=  $4\pi$  steradians**

# Hemispherical Coordinates

- Direction,  $\Theta$ 
  - Point on (unit) sphere



$$dA = (r \sin \theta d\varphi)(r d\theta)$$

From kavita's slides

# Hemispherical Coordinates

---

---

- **Differential solid angle**

$$d\omega = \frac{dA}{r^2} = \sin\theta d\theta d\varphi$$

# Hemispherical Integration

---

---

- Area of hemisphere:

$$\begin{aligned}\int_{\Omega_x} d\omega &= \int_0^{2\pi} d\varphi \int_0^{\pi/2} \sin\theta d\theta \\ &= \int_0^{2\pi} d\varphi [-\cos\theta]_0^{\pi/2} \\ &= \int_0^{2\pi} d\varphi \\ &= 2\pi\end{aligned}$$

# Energy

---

---

- **Symbol:  $Q$** 
  - **# of photons in this context**
  - **Unit: Joules**



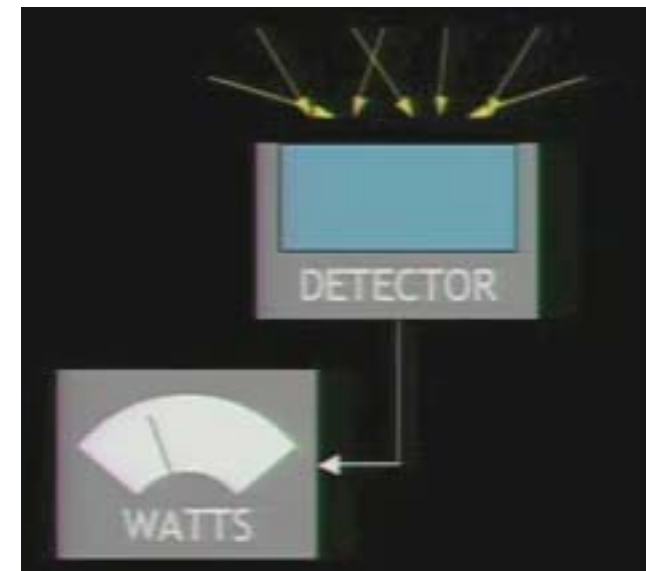
From Steve Marschner's talk

# Power (or Flux)

---

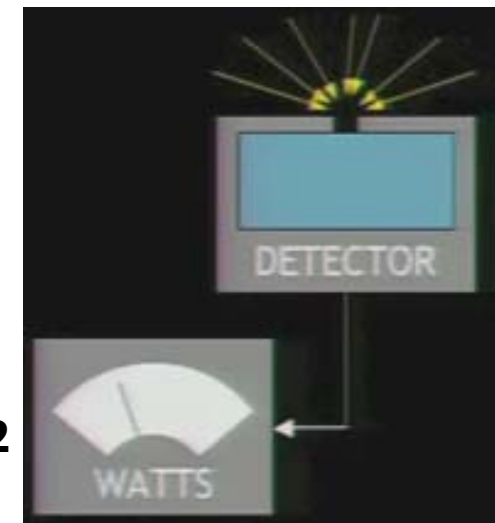
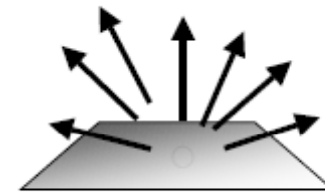
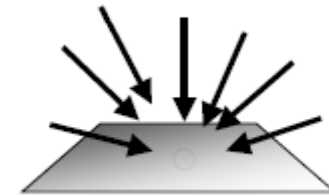
---

- **Symbol, P or  $\Phi$** 
  - Total amount of energy through a surface per unit time,  $dQ/dt$
  - Radiant flux in this context
  - Unit: Watts (=Joules / sec.)
  - Other quantities are derivatives of P
  
- **Example**
  - A light source emits 50 watts of radiant power
  - 20 watts of radiant power is incident on a table



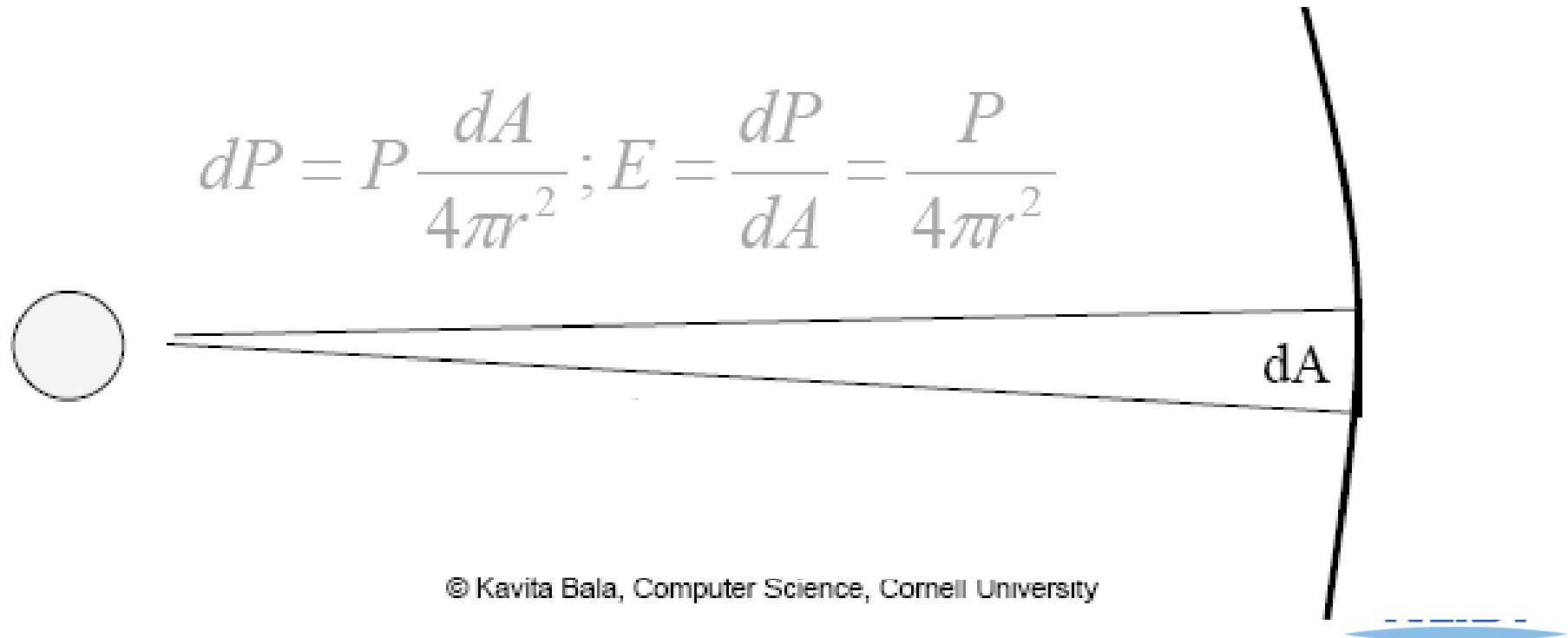
# Irradiance

- **Incident radiant power per unit area ( $dP/dA$ )**
  - Area density of power
- **Symbol:  $E$ , unit:  $W/m^2$** 
  - Area power density existing a surface is called radiance exitance (M) or radiosity (B)
- **For example**
  - A light source emitting 100 W of area  $0.1 m^2$
  - Its radiant exitance is  $1000 W/m^2$



# Irradiance Example

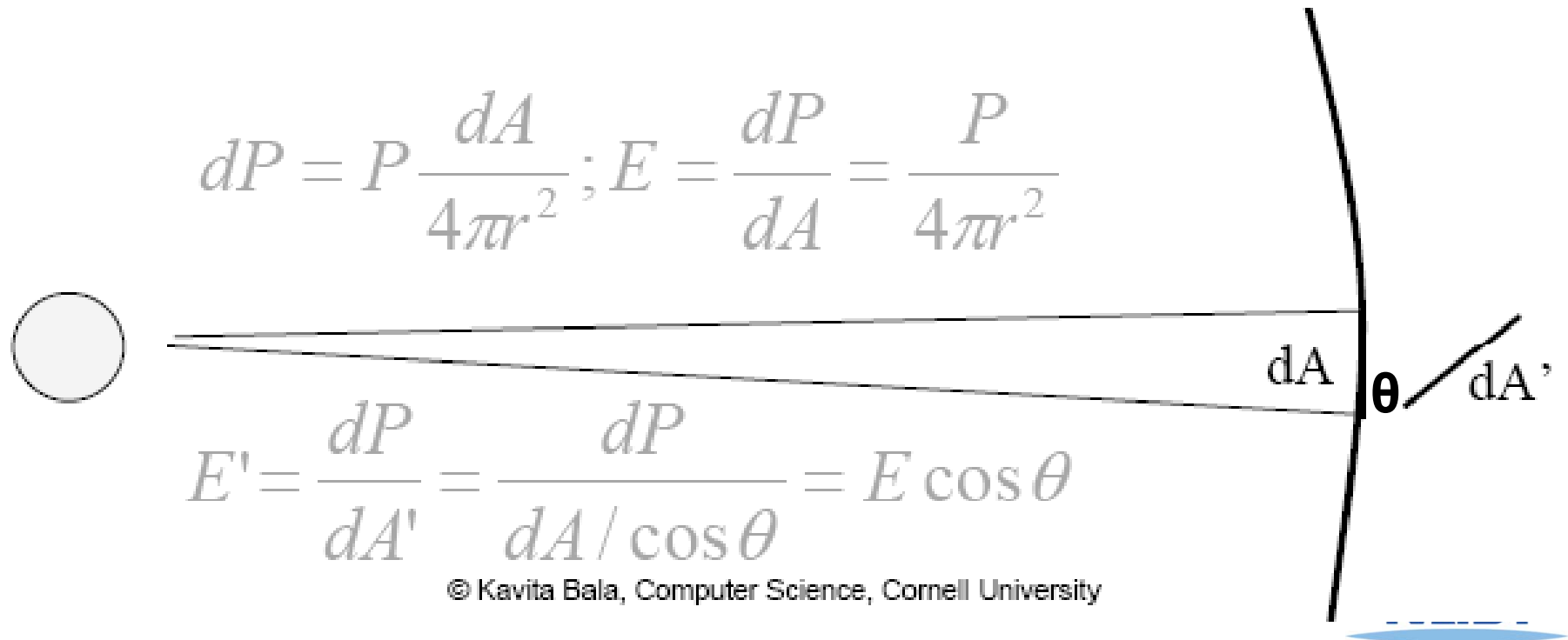
- **Uniform point source illuminates a small surface  $dA$  from a distance  $r$** 
  - **Power  $P$  is uniformly spread over the area of the sphere**





# Irradiance Example

- **Uniform point source illuminates a small surface  $dA$  from a distance  $r$** 
  - **Power  $P$  is uniformly spread over the area of the sphere**



# Radiance

---

---

- **Radiant power at  $x$  in direction  $\theta$** 
  - $L(x \rightarrow \Theta)$  : 5D function
    - Per unit area
    - Per unit solid angle

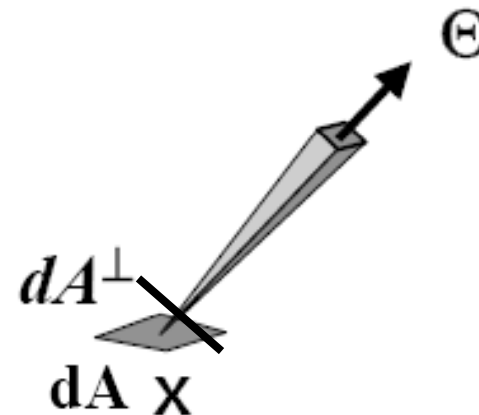


- **Important quantity for rendering**

# Radiance

- **Radiant power at  $x$  in direction  $\Theta$** 
  - $L(x \rightarrow \Theta)$  : 5D function
    - Per unit area
    - Per unit solid angle

$$L(x \rightarrow \Theta) = \frac{d^2 P}{dA^\perp d\omega_\Theta}$$

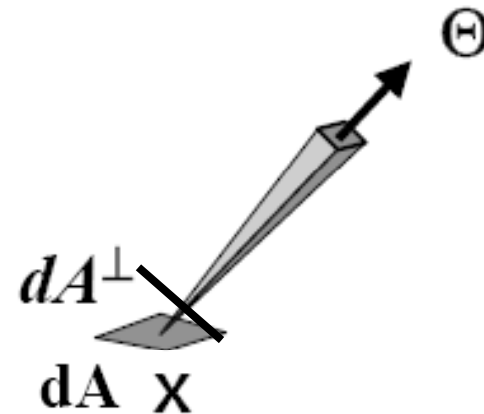


- **Units: Watt / (m<sup>2</sup> sr)**
- **Irradiance per unit solid angle**
- **2<sup>nd</sup> derivative of P**
- **Most commonly used term**

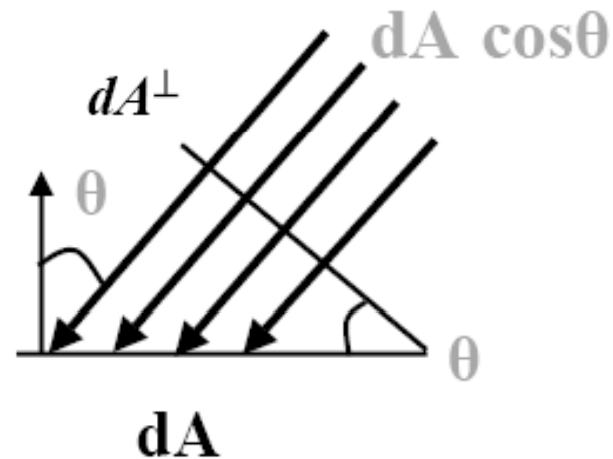
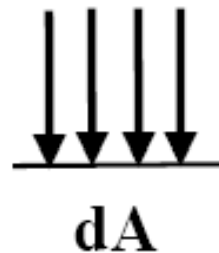
# Radiance: Projected Area

$$L(x \rightarrow \Theta) = \frac{d^2 P}{dA^\perp d\omega_\Theta}$$

$$= \frac{d^2 P}{d\omega_\Theta dA \cos \theta}$$

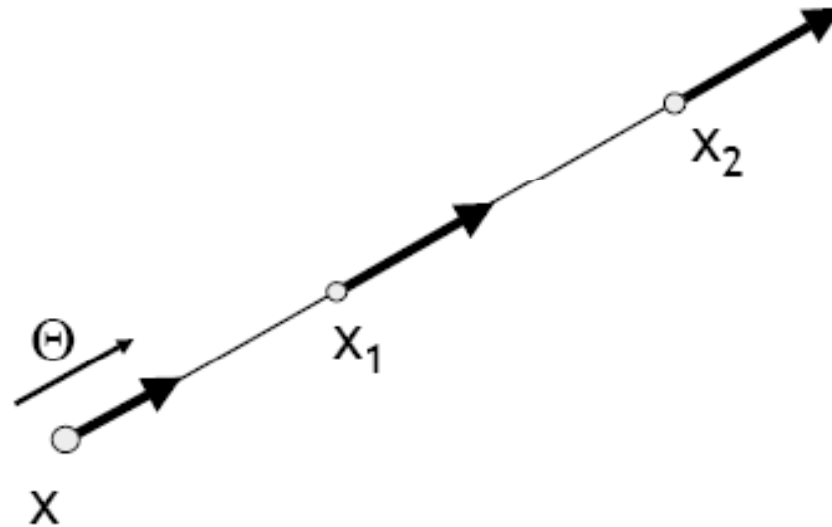


- Why per unit projected surface area



# Properties of Radiance

- **Invariant along a straight line (in vacuum)**



From kavita's slides

# Invariance of Radiance

We can prove it based on the assumption the conservation of energy.

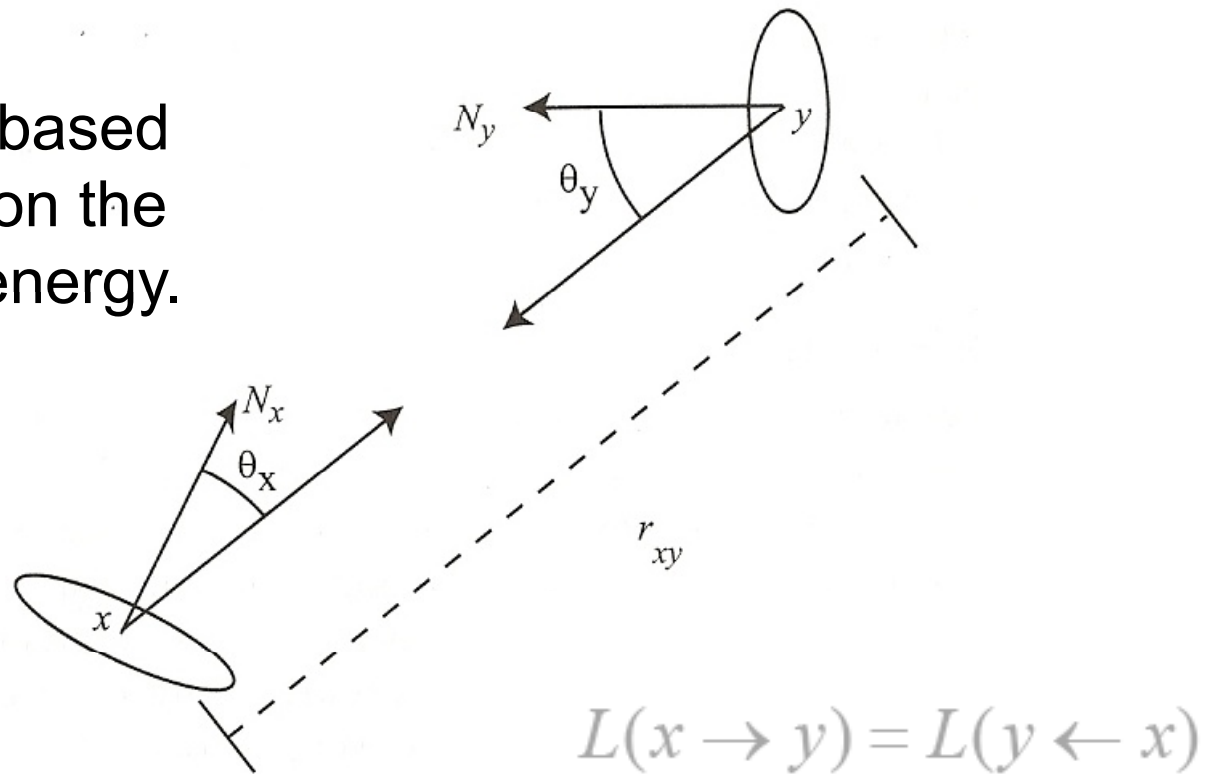


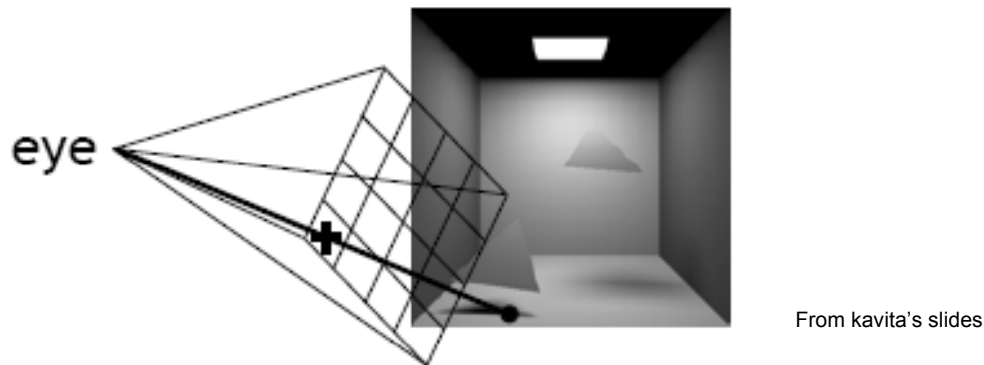
Figure 2.3. Invariance of radiance.

# Sensitivity to Radiance

---

---

- **Responses of sensors (camera, human eye) is proportional to radiance**



- **Pixel values in image proportional to radiance received from that direction**

# Relationships

---

---

- Radiance is the fundamental quantity

$$L(x \rightarrow \Theta) = \frac{d^2P}{dA^\perp d\omega_\Theta}$$

- Power:

$$P = \int_{\text{Area}} \int_{\text{Solid Angle}} L(x \rightarrow \Theta) \cdot \cos \theta \cdot d\omega_\Theta \cdot dA$$

- Radiosity:

$$B = \int_{\text{Solid Angle}} L(x \rightarrow \Theta) \cdot \cos \theta \cdot d\omega_\Theta$$

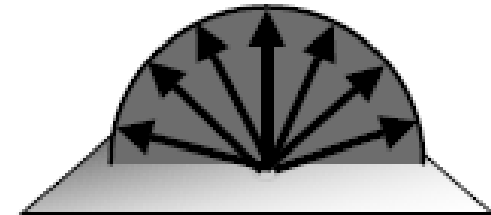


# Example: Diffuse emitter

---

- Diffuse emitter: light source with equal radiance everywhere

$$L(x \rightarrow \Theta) = \frac{d^2 P}{dA^\perp d\omega_\Theta}$$

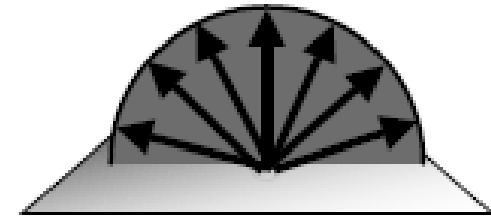


# Example: Diffuse emitter

---

- Diffuse emitter: light source with equal radiance everywhere

$$L(x \rightarrow \Theta) = \frac{d^2 P}{dA^\perp d\omega_\Theta}$$



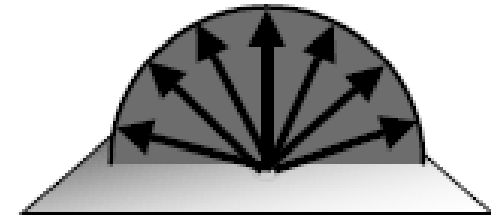
$$P = \int_{\text{Area}} \int_{\text{Solid Angle}} L(x \rightarrow \Theta) \cdot \cos \theta \cdot d\omega_\Theta \cdot dA$$

# Example: Diffuse emitter

---

- Diffuse emitter: light source with equal radiance everywhere

$$L(x \rightarrow \Theta) = \frac{d^2 P}{dA^\perp d\omega_\Theta}$$

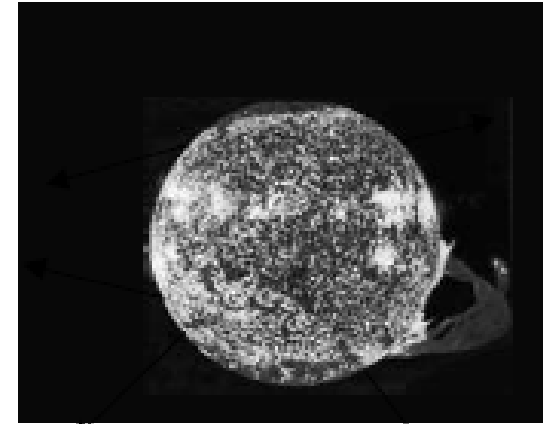


$$\begin{aligned} P &= \int_{\substack{\text{Area} \\ \text{Angle}}} \int_{\substack{\text{Solid} \\ \text{Angle}}} L(x \rightarrow \Theta) \cdot \cos \theta \cdot d\omega_\Theta \cdot dA \\ &= L \int_{\substack{\text{Area} \\ \text{Area}}} dA \int_{\substack{\text{Solid} \\ \text{Angle}}} \cos \theta \cdot d\omega_\Theta \\ &= L \cdot \text{Area} \cdot \pi \end{aligned}$$

# Sun Example: radiance

---

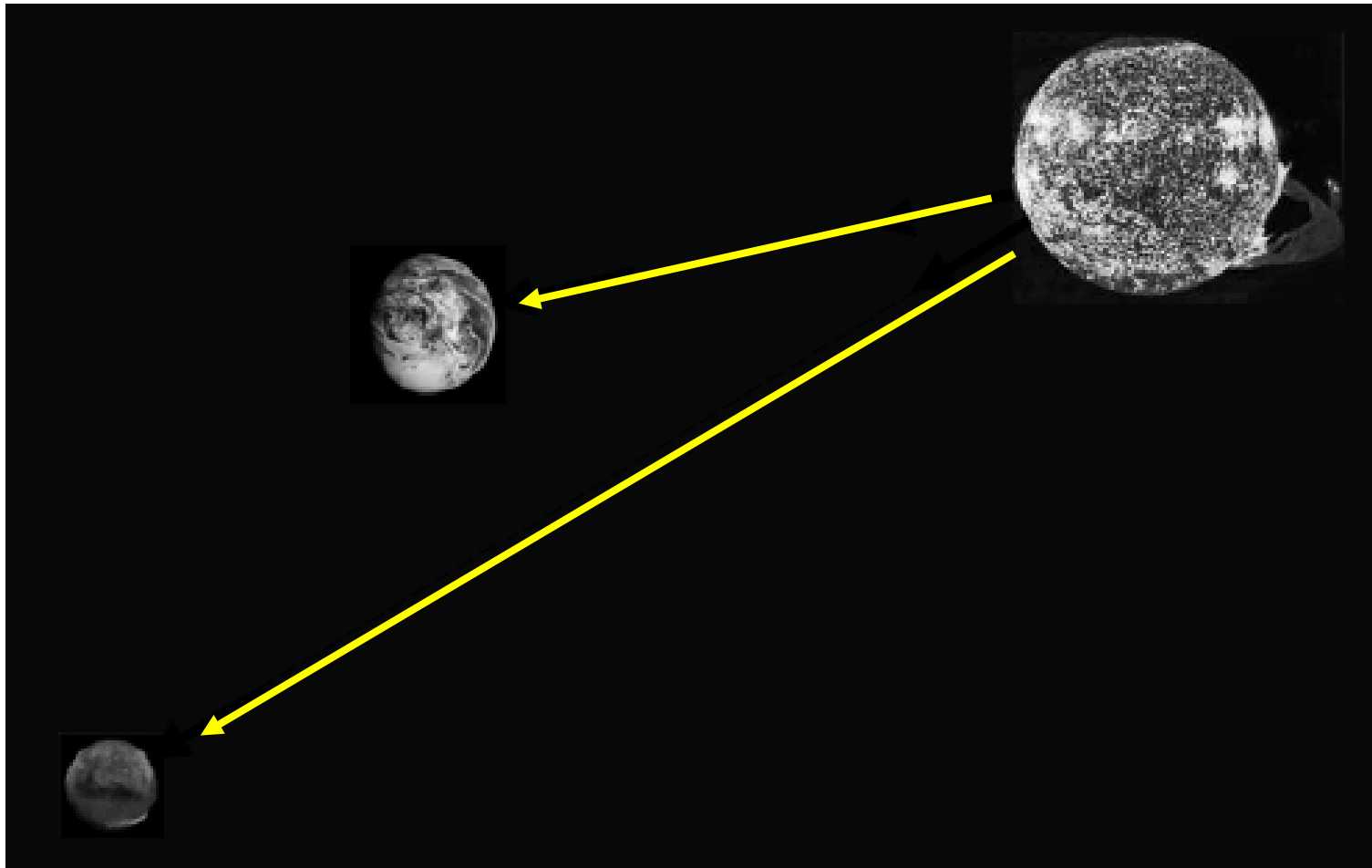
- Power:  $3.91 \times 10^{26} \text{ W}$
- Surface Area:  $6.07 \times 10^{18} \text{ m}^2$



- Power = Radiance  $\cdot$  Surface Area  $\cdot \pi$
- Radiance = Power / (Surface Area  $\cdot \pi$ )
- Radiance =  $2.05 \times 10^7 \text{ W/ m}^2 \cdot \text{sr}$

# Sun Example

---



**Same radiance on Earth and Mars?**

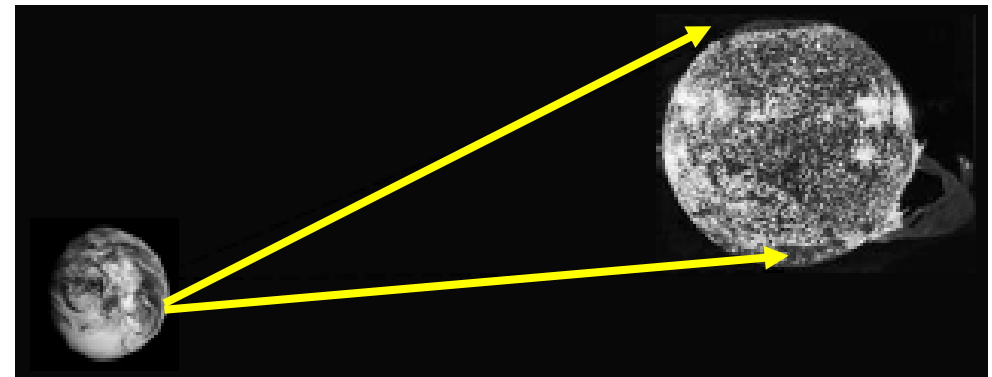
© Kavita Bala, Computer Science, Cornell University

# Sun Example: Power on Earth

---

- Power reaching earth on a 1m<sup>2</sup> square:

$$P = L \int_{\text{Area}} dA \int_{\text{Solid Angle}} \cos\theta \cdot d\omega_{\odot}$$



- Assume  $\cos\theta = 1$  (sun in zenith)

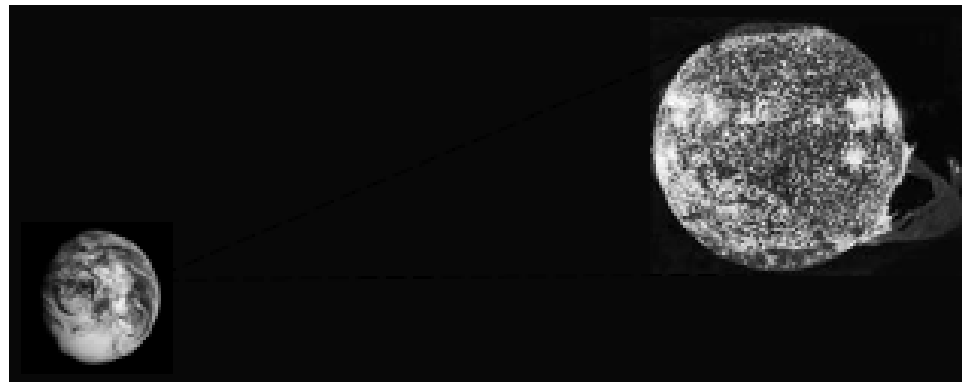
$$P = L \int_{\text{Area}} dA \int_{\text{Solid Angle}} d\omega_{\odot}$$

---

# Sun Example: Power on Earth

---

**Power = Radiance.Area.Solid Angle**



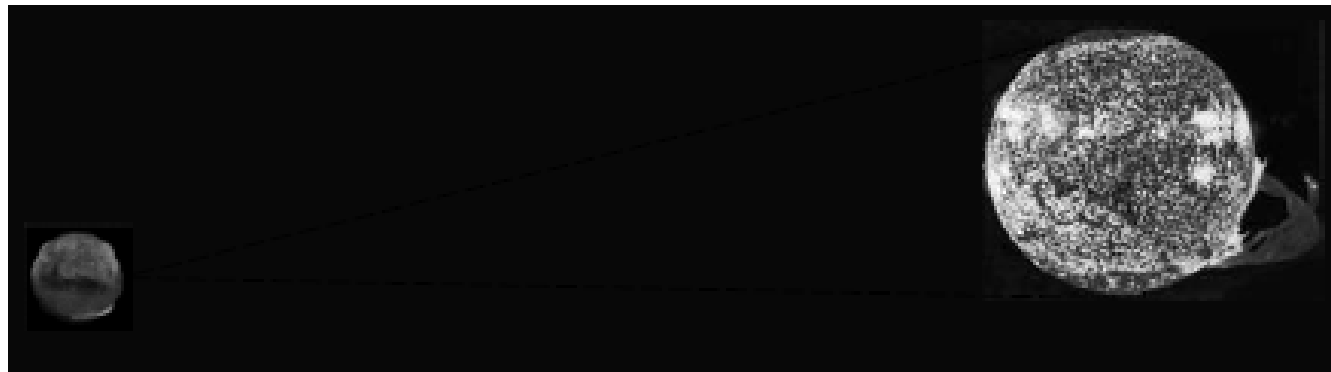
$$\begin{aligned}\text{Solid Angle} &= \text{Projected Area}_{\text{Sun}} / (\text{distance}_{\text{earth\_sun}})^2 \\ &= 6.7 \cdot 10^{-5} \text{ sr}\end{aligned}$$

$$\begin{aligned}P &= (2.05 \times 10^7 \text{ W/ m}^2\cdot\text{sr}) \times (1 \text{ m}^2) \times (6.7 \cdot 10^{-5} \text{ sr}) \\ &= 1373.5 \text{ Watt}\end{aligned}$$

# Sun Example: Power on Mars

---

**Power = Radiance.Area.Solid Angle**



$$\begin{aligned}\text{Solid Angle} &= \text{Projected Area}_{\text{Sun}} / (\text{distance}_{\text{mars\_sun}})^2 \\ &= 2.92 \cdot 10^{-5} \text{ sr}\end{aligned}$$

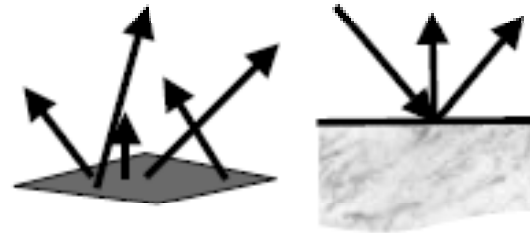
$$\begin{aligned}P &= (2.05 \times 10^7 \text{ W/ m}^2 \cdot \text{sr}) \times (1 \text{ m}^2) \times (2.92 \cdot 10^{-5} \text{ sr}) \\ &= 598.6 \text{ Watt}\end{aligned}$$



# Light and Material Interactions

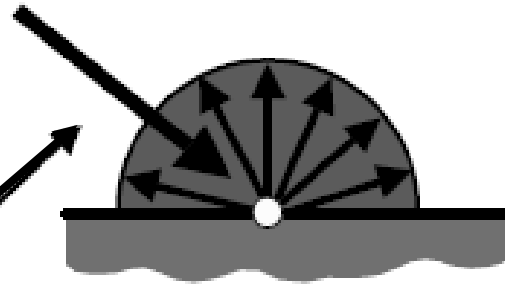
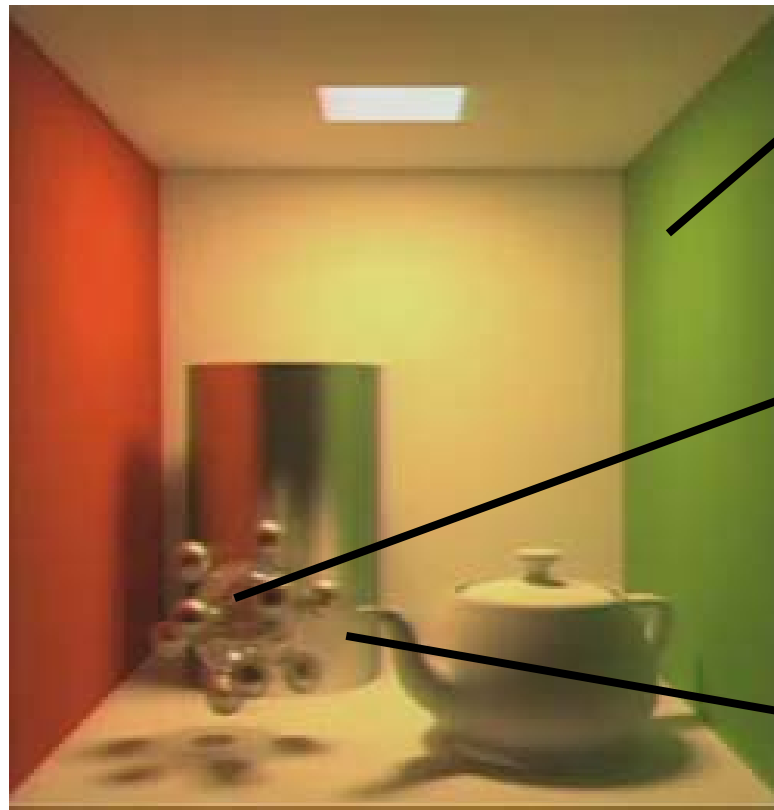
---

- Physics of light
- Radiometry
- **Material properties**
- Rendering equation



From kavita's slides

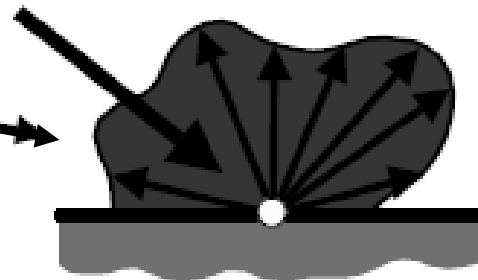
# Materials



**Ideal diffuse  
(Lambertian)**



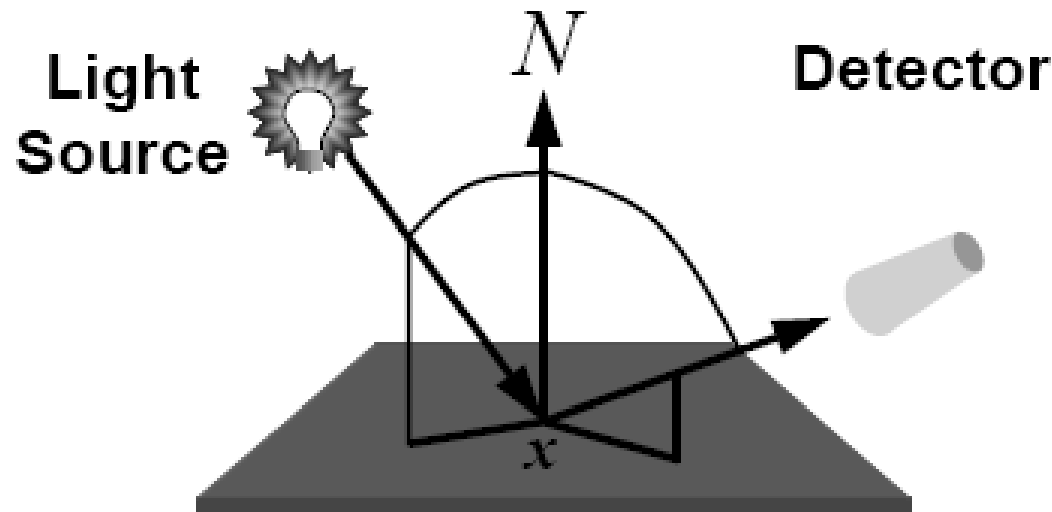
**Ideal specular**



**Glossy**

From kavita's slides

# Bidirectional Reflectance Distribution Function (BRDF)



$$f_r(x, \Psi \rightarrow \Theta) = \frac{dL(x \rightarrow \Theta)}{dE(x \leftarrow \Psi)} = \frac{dL(x \rightarrow \Theta)}{L(x \leftarrow \Psi) \cos(N_x, \Psi) d\omega_\Psi}$$

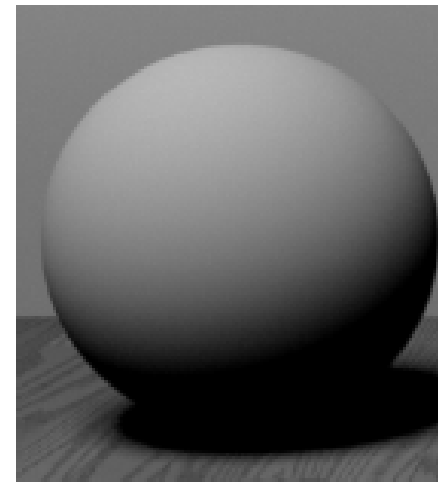
© Kavita Bala, Computer Science, Cornell University

# BRDF special case: ideal diffuse

---

## Pure Lambertian

$$f_r(x, \Psi \rightarrow \Theta) = \frac{\rho_d}{\pi}$$



$$\rho_d = \frac{\text{Energy}_{out}}{\text{Energy}_{in}} \quad 0 \leq \rho_d \leq 1$$

# Properties of the BRDF

---

- Reciprocity:

$$f_r(x, \Psi \rightarrow \Theta) = f_r(x, \Theta \rightarrow \Psi)$$

- Therefore, notation:  $f_r(x, \Psi \leftrightarrow \Theta)$
- Important for bidirectional tracing

# Properties of the BRDF

---

- Bounds:

$$0 \leq f_r(x, \Psi \leftrightarrow \Theta) \leq \infty$$

- Energy conservation:

$$\forall \Psi \int_{\Theta} f_r(x, \Psi \leftrightarrow \Theta) \cos(N_x, \Theta) d\omega_{\Theta} \leq 1$$

# Homework

---

---

# Next Time

---

---

- **Rendering equation**